

## Lesson 15: Real-World Development of a Lens

In Lesson 14 we designed a 7-element lens starting with nothing but plane-parallel surfaces and had the program fit the design to catalog glass types automatically with the ARGLASS feature. But suppose you have a real application and want to develop it further. This lesson covers some additional procedures that would then be appropriate. To make it a real “real world” lesson, we will show how a designer will follow various clues in order to arrive at a solution, and how not all clues lead to success. That is important too: it is instructive to see how sometimes one wanders into blind alleys. As you develop your skills as a lens designer, you will encounter many of them, and should not be discouraged since it happens to us all. With perseverance, a successful design can usually be found.

We will do this lesson in two ways; first with DSEARCH with the help of a number of other tools. Then, in Lesson 17 we show another approach that is actually quicker and easier. You should know about all of the tools used in both approaches.

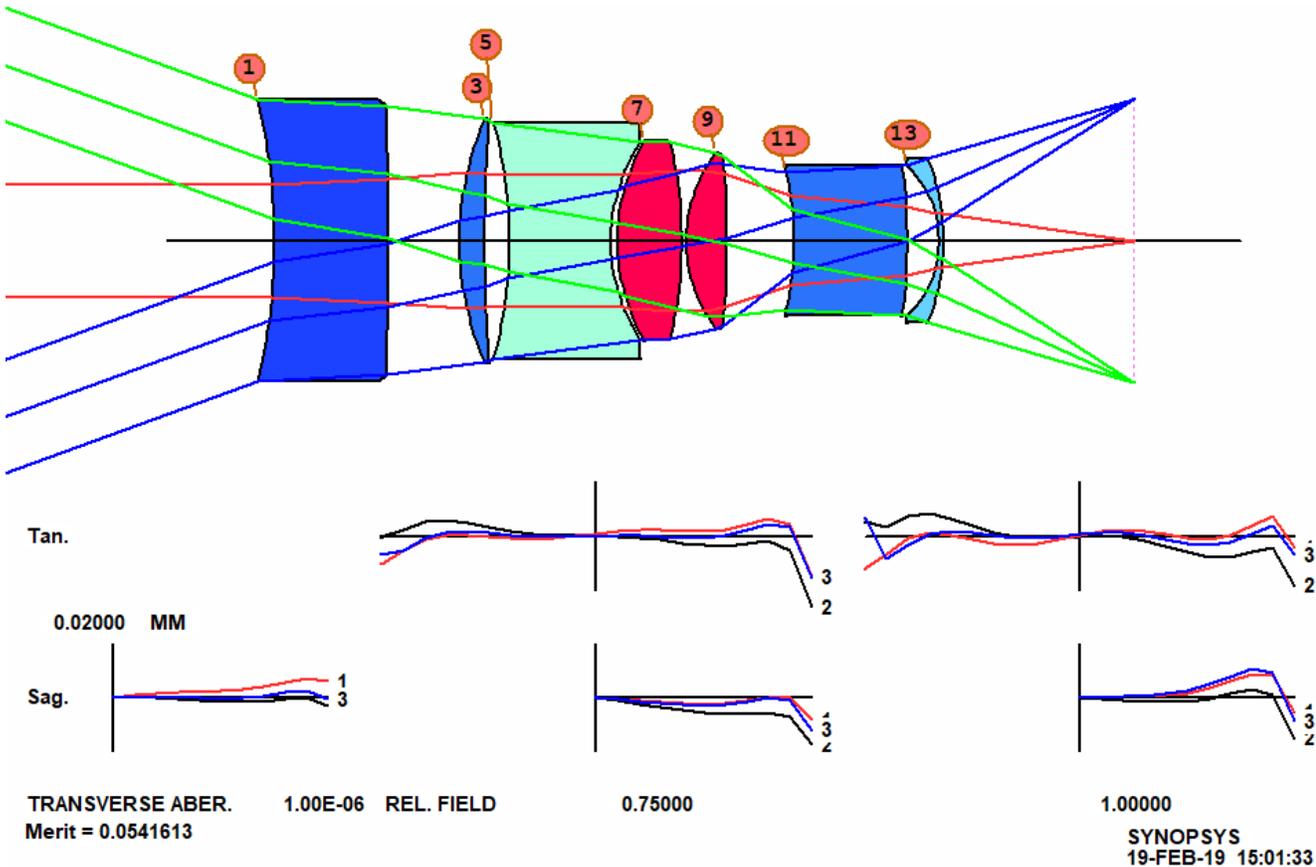
We will first use DSEARCH to find a good starting point. Here is the input:

```
CORE 14
DSEARCH 6 QUIET
SYSTEM
ID DSEARCH SAMPLE
OBB 0 20 12.7
WAVL 0.6563 0.5876 0.4861

UNITS MM
END
GOALS
ELEMENTS 7
FNUM 3.575
BACK 50 .01
STOP MIDDLE
STOP FREE
RT 0.5
FOV 0.0 0.75 1.0 0.0 0.0
FWT 5.0 3.0 3
DELAY 999
RSTART 900
THSTART 7
ASTART 15
NPASS 66
ANNEAL 200 20 Q
COLORS 3
SNAPSHOT 10
QUICK 44 66
END
SPECIAL PANT

END
SPECIAL AANT
LUL 150 1 1 A TOTL
END
GO
```

We run this, and the best lens that is returned is quite good, except for thin elements, which we will fix up later. We optimize and anneal, using the file DSEARCH\_OPT, which is in a new editor window.



Suppose we want the lens to work over a range of object distances from one meter to infinity. There are two ways to implement that requirement: with multiconfigurations, which is very flexible but complicated, or by declaring this a zoom lens in which the object distance zooms. The second approach is better here, since it is simpler, does what we want, and we can examine intermediate object distances very easily. We have to set up this lens as a ZFILE zoom lens.

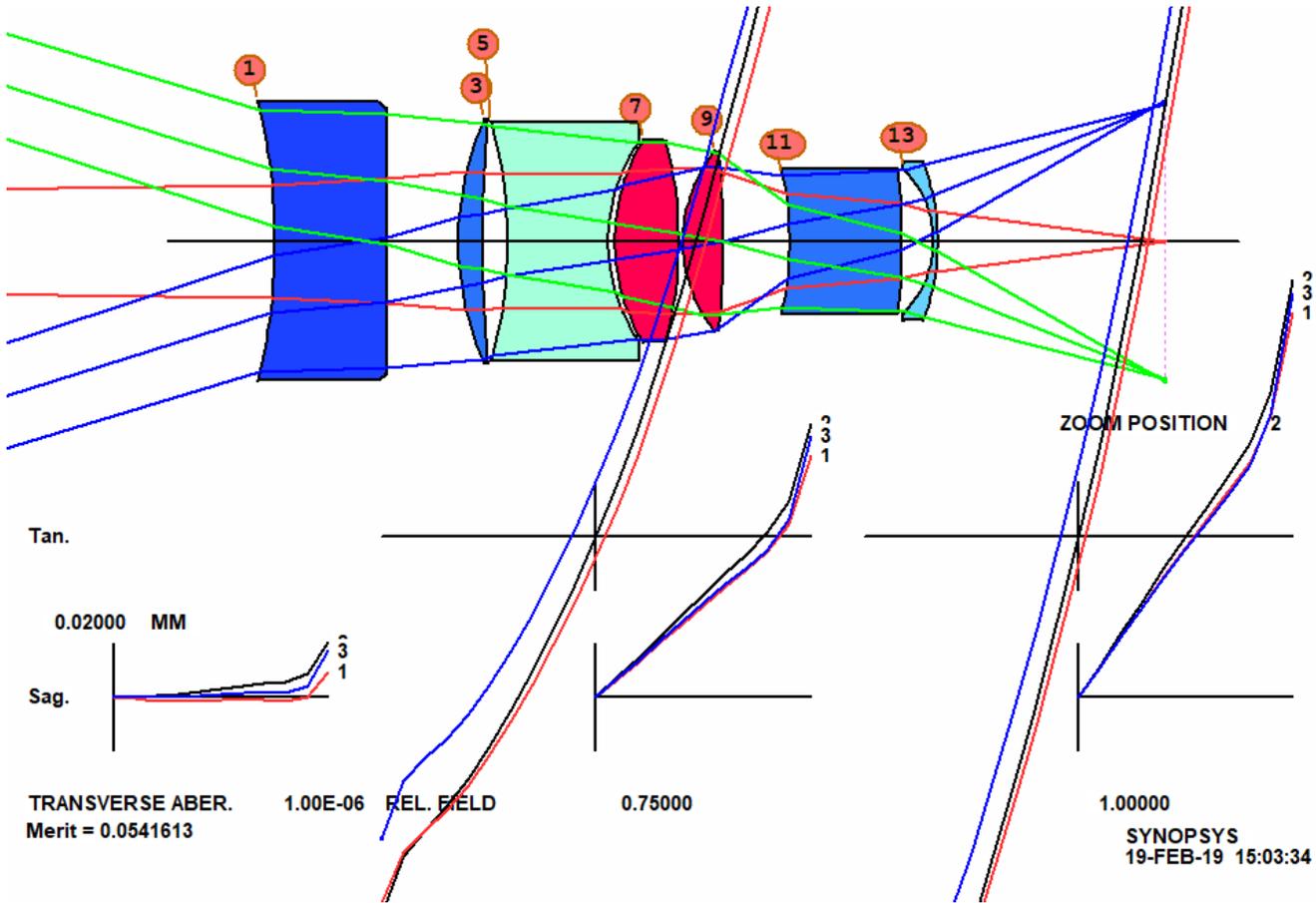
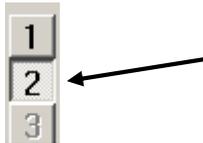
```

CHG
APS -10           ! declare surface 10 the stop
15 CAO 32         ! fix the CAO on the image (so FFIELD works)
FFIELD           ! adjust the object height so the image fills the CAO there
14 YMT           ! assign a paraxial focus solve to surface 14
ZFILE 1          ! start of the ZFILE section
14 14            ! there is one zooming group, the last thickness
ZOOM 2           ! ZOOM 1 is default; ZOOM 2 gets OBA object on the next line
OBA 1000 -366.554 12.7 ! the object description at this zoom
END              ! end of changes

```

Here we declare surface 10 a real stop, so all zooms use the same location, set a hard aperture at the image so the FFIELD directive has a target, put a thickness solve on 14 so all zooms refocus automatically, and declare a single zooming group, surface 14. Then we define the object distance for ZOOM 2 at 1000 mm distance, with a negative YPPO because the value in ZOOM 1 is also negative, and they have to have the same sign.

Run this MACro, and the lens changes to a zoom lens, with only a single airspace zooming in this case. Now you see a new toolbar on the right side of the monitor. What does the image look like in ZOOM 2? If you click on buttons 1 and 2 you see the lens at that zoom setting. Here is zoom 2:



Pretty awful! We have to correct the image at both conjugates. Here is our MACro:

```

AWT: 0.5
PANT      ! Define variables.
!VY 1 YP1 ! Don't vary YP1; it is not compatible with the real pupil declaration
VLIST RAD ALL      ! Varies all radii that are not flat.
VLIST TH ALL       ! varies all thicknesses and airspaces except for the
! back focus, thickness 14, which has a solve in effect
VLIST GLM ALL
END

AANT      ! Start of merit function definition.
AEC       ! Activate automatic edge-feathering monitor
ACC       ! and maximum center thickness monitor.
ADT 6 .1 10      ! Keep diameter/thickness ratio 6 or more
!M 33 2 A GIHT   ! Comment this out, since the FFIELD will control scale
LUL 150 1 1 A TOTL
M 50 .1 A BACK   ! Since the back focus will vary, keep it reasonable
M 90.61 1 A FOCL ! Add this requirement so the focal length doesn't change
GSR AWT 10 5 M 0 ! Note how weights are assigned to the several field points,
! and the symbol AWT controls the aperture weighting.
GNR AWT 5.5 4 M .5      ! This creates a ray grid at the 1/2 field point
GNR AWT 5.5 4 M .7      ! These for the 0.7 field point
GNR AWT 3 4 M 1        ! Full field gets the lowest weight.

```

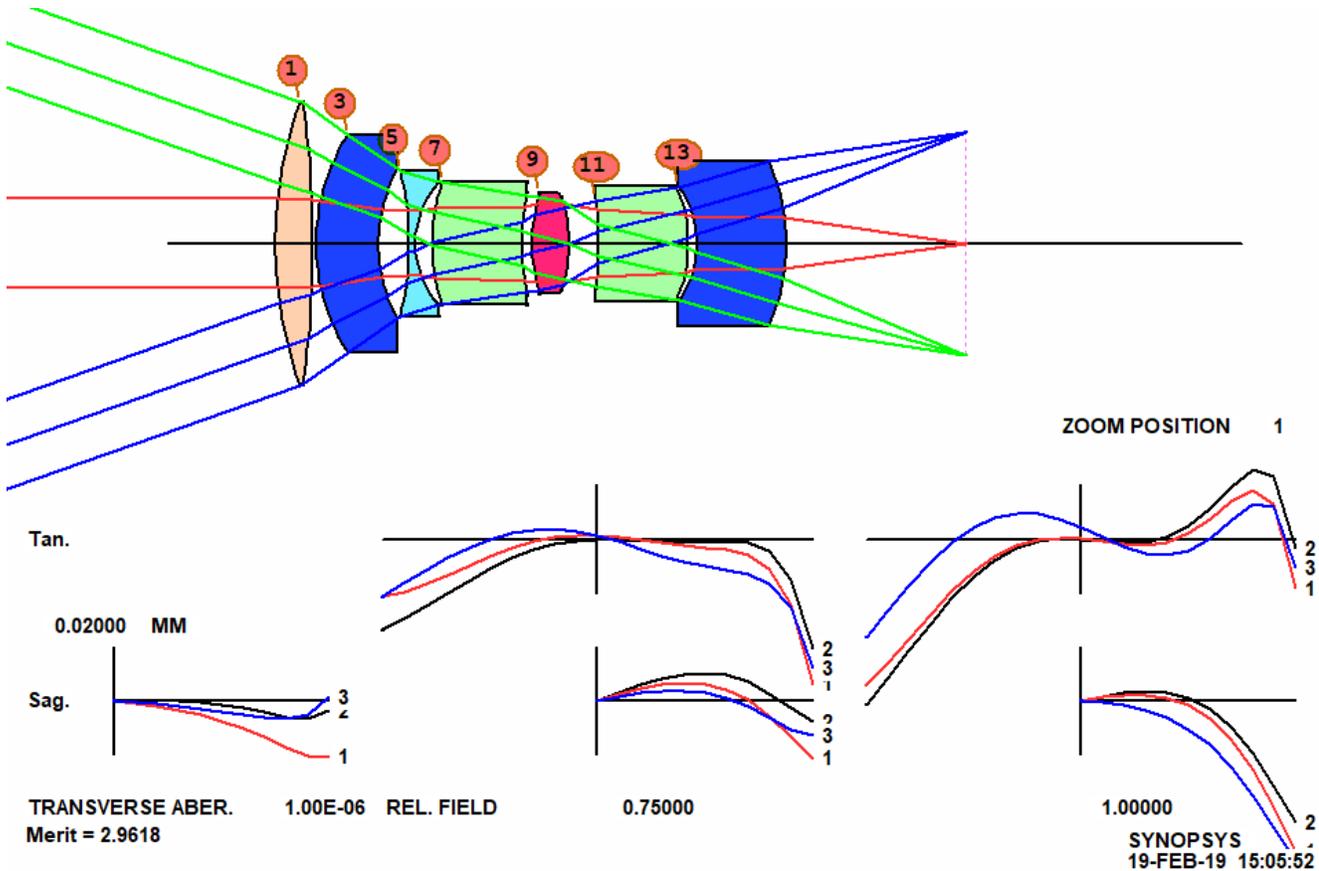
```

ZOOM 2          ! Targets for zoom 2 (with the object at one meter)
GSR AWT 10 5 M 0 ! Note how weights are assigned to field points.
GNR AWT 5.5 4 M .5 ! This creates a ray grid at the 1/2 field point
GNR AWT 5.5 4 M .7 ! These for the 0.7 field point
GNR AWT 3 4 M 1    ! Full field gets the lowest weight.
END

SNAP
SYNO 50

```

Run this and anneal, and the lens is better but still not very good, with about equal errors at both ends of the zoom range.

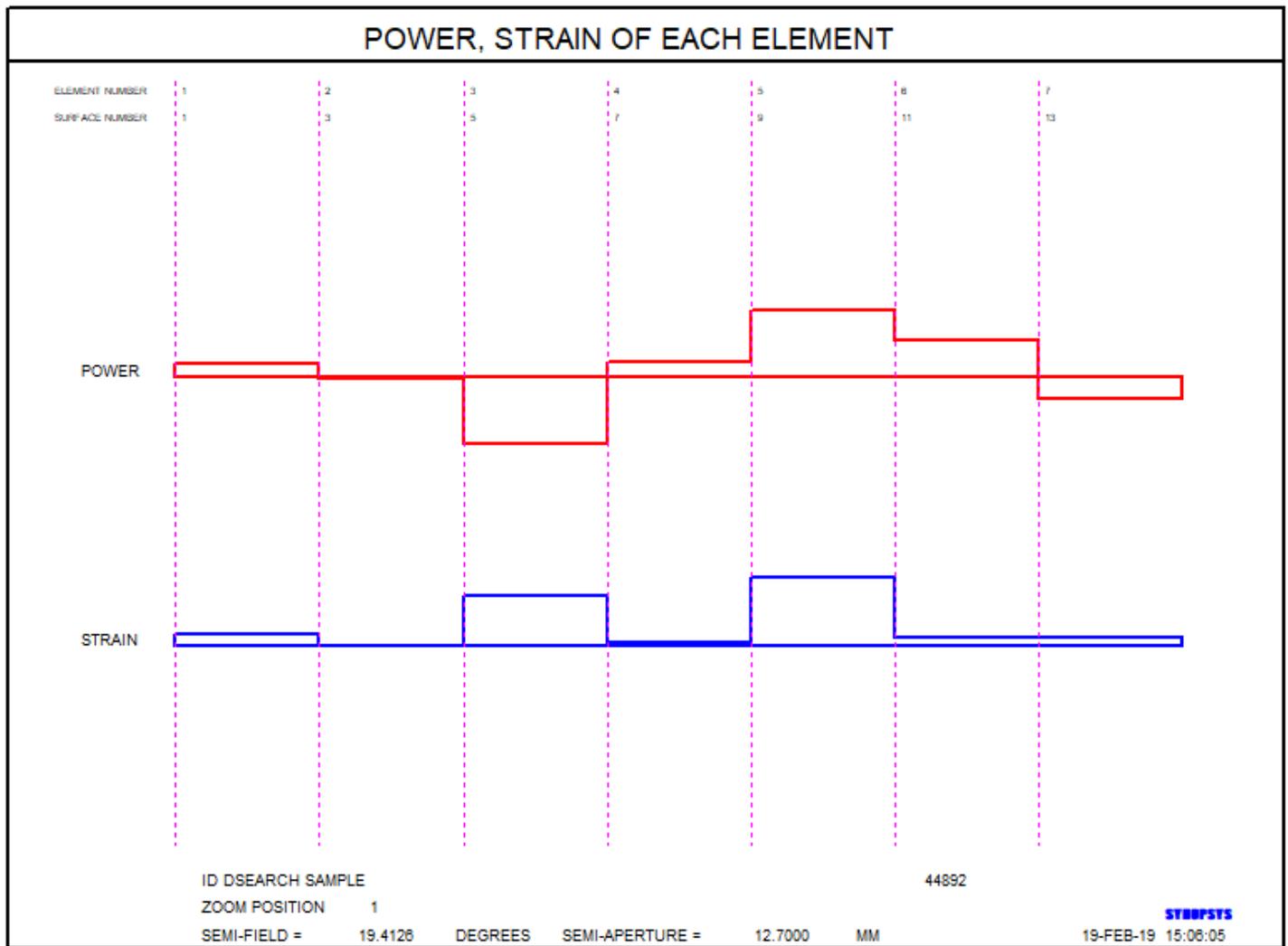


Some subtleties deserve mention: the **GLM ALL** variable will vary all glass models currently in the lens, which means all elements, since DSEARCH uses the glass model unless told otherwise. We have to control the focal length since the object height will be continuously adjusted so the image CAO is filled at full field.

This is better than zoom 2 was before, but there is still a loss of resolution. What to do? We need more variables. What should we add?

A classic tool for cases like this is the STRAIN calculation. The idea is that the surfaces with the largest strain are contributing most of the low-order aberrations, and splitting an element there might relieve that strain.

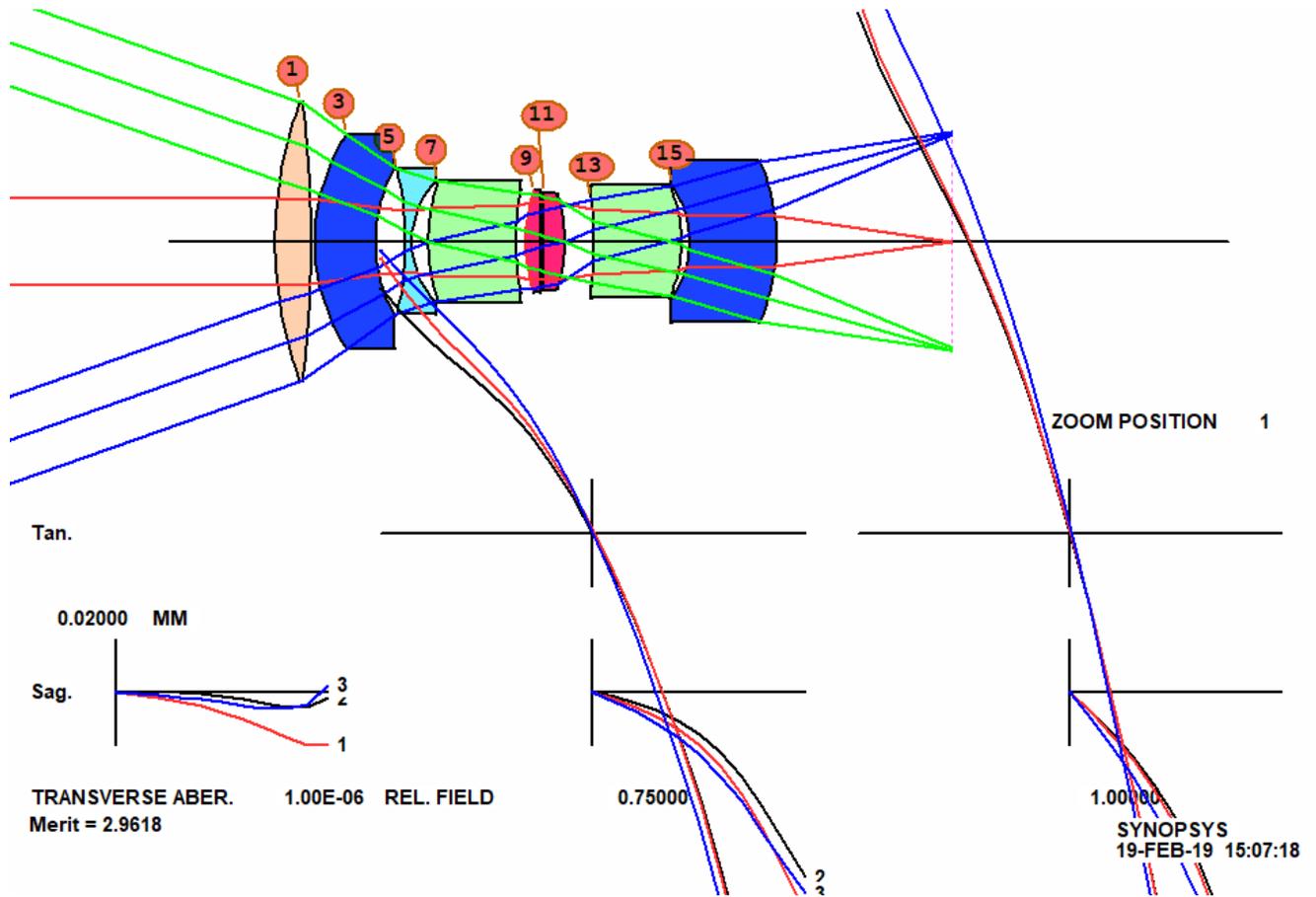
Type **STRAIN P** in the CW.



Indeed, element five has the largest strain. Now we can do one of two things: We can split that element and reoptimize, or we can use a different tool that can figure out the best place to add an element. We will try it both ways. First, let's save this version, so we can go back if things don't work out. Type

STORE 1.

Then go to the WorkSheet (type WS, or click on the button ). Then click the button , which lets you split an element by clicking in the PAD display on the axis inside that element. Click between surfaces 9 and 10, splitting the element. Your lens now looks like this:



When the program splits (or adds) an element, it assigns an index pickup, because at that moment it has no other index data. In WS, change the index pickup on surface 11 to a glass model by typing

**11 GLM**

in the edit pane, and click Update. That changes to a model glass with properties similar to what were there before.

Make a new checkpoint, close WS, run the optimization again, and we see that the lens has improved slightly. The MF is now 1.59. This is the way lens design has long been done, using classic tools, and it was a slow and arduous process. But today we have better tools. Go back to the version before you split the element:

**GET 1**

and then add a line before the PANT file:

**AEI 2 1 14 0 0 0 10 2**

This will run the Automatic Element Insertion tool (**AEI**). Now the program will search for the best place to insert a new element. Run this, and the lens is better. Comment out the AEI line and run your MACro again, then anneal. Here is the result:

```
RLE
ID DSEARCH SAMPLE                44892
ID1 DSEARCH CASE WAS 00000000000000000000100011    35
FNAME 'AEI08.RLE
MERIT  1.49487
LOG    44892
```

WAVL .6563000 .5876000 .4861000

APS -12

FFIELD

UNITS MM

OBB	0.000000	19.41264	12.70000	-27.20408	0.00000	0.00000	12.70000
0	AIR						
1	RAD	126.2610314937226	TH	5.65004368			
1	GLM	1.71002696		57.03505884			
2	RAD	548.9829634420437	TH	1.00000000	AIR		
3	RAD	48.8001314259707	TH	7.15931650			
3	GLM	1.90000000		30.45165541			
4	RAD	78.8914205169275	TH	3.92580954	AIR		
5	RAD	144.0284949074012	TH	4.36091211			
5	GLM	1.58044211		40.30348968			
6	RAD	22.8152407802378	TH	8.52017600	AIR		
7	RAD	41.6567444491098	TH	13.58034863			
7	GLM	1.90000000		29.97777667			
8	RAD	24.7675453690692	TH	14.60211862	AIR		
9	RAD	80.8352977692681	TH	10.27627981			
9	GLM	1.57399779		41.33289870			
10	RAD	43.3646405210182	TH	1.29508471	AIR		
11	RAD	49.7804194802432	TH	10.66304480			
11	GLM	1.50100000		77.45538462			
12	RAD	-28.4841065971860	TH	39.36991857	AIR		
13	RAD	-337.5623606237147	TH	4.69577023			
13	GLM	1.64849716		63.04604701			
14	RAD	-53.4480004366635	TH	3.02049855	AIR		
15	RAD	-38.5862175963575	TH	22.01446071			
15	GLM	1.56813982		42.33714495			
16	RAD	-74.9651540676835	TH	50.65153648	AIR		
16	CV	-0.01333953					
16	UMC	-0.13986014					
16	TH	50.65153648					
16	YMT	0.00000000					
17	CAO	32.00000000	0.00000000	0.00000000			
17	CV	0.00000000000000	TH	0.00000000	AIR		

ZFILE 1

CAM RANK 2

CAM EXPONENT 1.00000

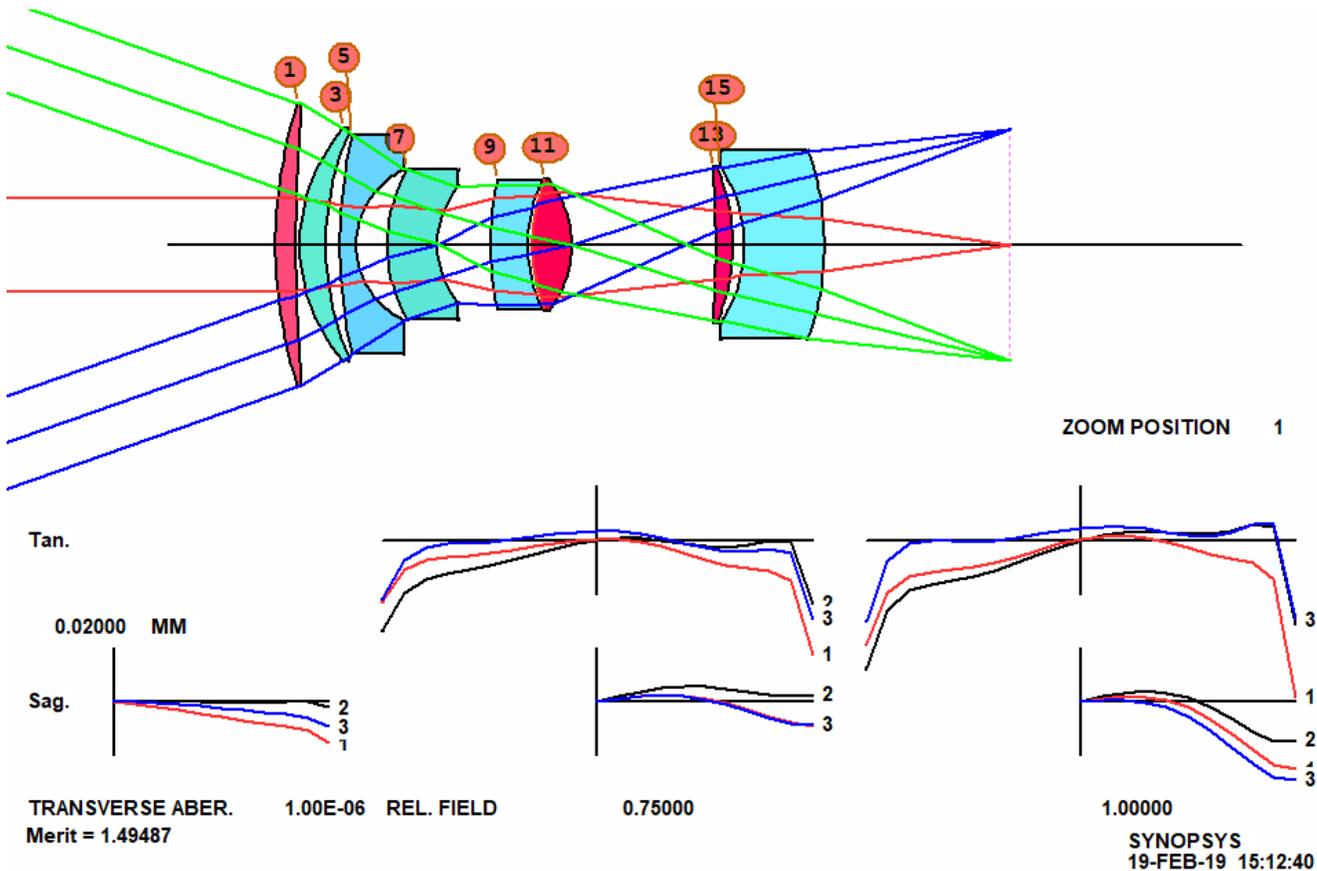
16 16

ZOOM 2

OBA	1000.00000	-366.554000	12.7000	0.0000	0.00000000	0.0000	12.7000
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ZDATA 0.0000000E+00

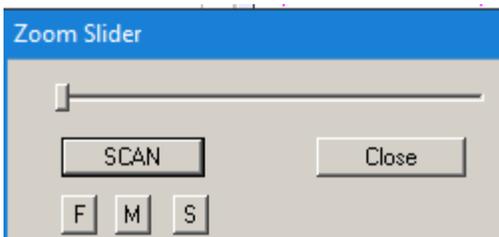
END



Wow! The program has inserted a new element at surface 9! And the merit function came down from about 2.96 to 1.49. There's a lesson here: The program can figure out how to improve a lens better than you can (unless you are very gifted). So it's better to let AEI do it than to try things that seem to make sense. Those things sometimes work, but AEI is better.

Here you see a larger improvement, and the MTF is also better, as you can check yourself. Now we have a lens that is fairly well corrected for both infinity conjugate and at one meter. But what about in-between distances? It would be a rude surprise if we built the lens and found that at an intermediate distance things got really bad. We have to check.

That is one of the reasons we chose to use the ZFILE zoom feature for this job. We can easily scan over the zoom range and spot any points that perhaps need attention. Click the button at the bottom of the zoom-selection bar:  This opens a zoom slider that is fun to watch.



Slide the thumb slowly to the right end, watching the PAD display (or click the SCAN button). The image plane slowly moves back, from the infinity focus to the one-meter focus position. The good news is, the image quality shows little change over the entire range, and in fact gets better near the middle. (If it had changed, we could use the CAM

command to create an intermediate focus position, making a total of three zooms, and then add some more targets for the ZOOM 3 position in the AANT file.) You can create and target up to 20 zooms, as you will learn if you type HELP CAM to read about that feature.

So we have roughed out a lens that works quite well over the entire focus range. Of course we are not yet done. Now we need to assign real glasses again, and it would be a good idea to increase the thickness of some of the elements, delete those thickness variables, and reoptimize. But ... wait a minute. The fourth element shown in the picture above bothers us. What is it doing? Use the STRAIN command again, and you see that there is very little power or strain on that element. That's a sign that we just might be able to remove it entirely. We have to try! Remove the AEI directive and replace it with

**AED 5 QUIET 1 15**

And run it again, and – wow! The program says that element at surface 9 can be removed. Allow it to do this, then comment out the AED directive and optimize some more. The merit function goes to 1.79 – not as good as before, but perhaps good enough. And we have eliminated an element. See how AED can make excellent decisions?

```

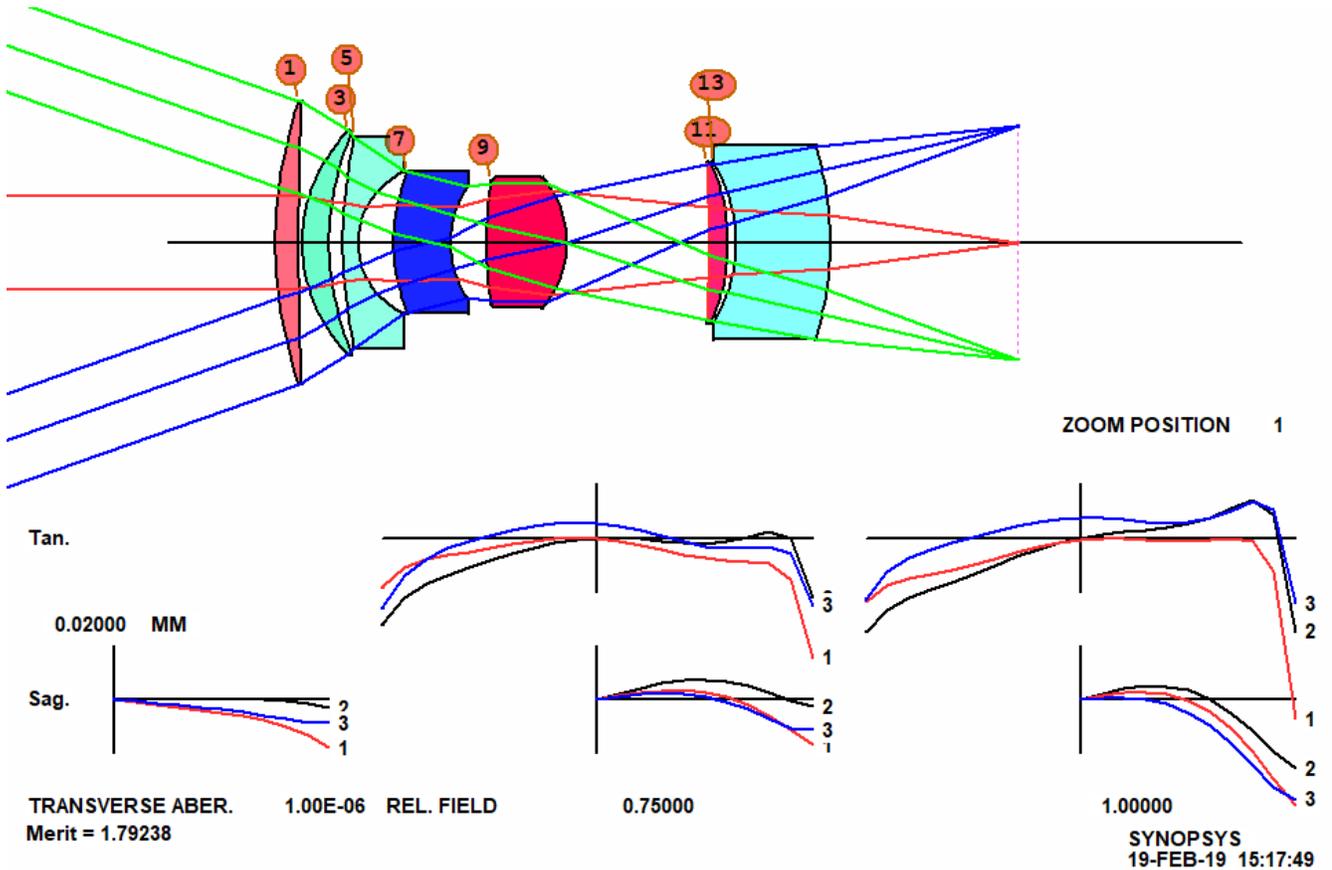
RLE
ID DSEARCH SAMPLE                44892
ID1 DSEARCH CASE WAS 00000000000000000100011    35
FNAME 'AEI08.RLE
MERIT 1.79238
LOG 44892
WAVL .6563000 .5876000 .4861000
APS -10
FFIELD
UNITS MM
OBB 0.000000    19.41264    12.70000    -26.64355    0.00000    0.00000    12.70000
0 AIR
1 RAD 120.8868431760548 TH 6.29742193
1 GLM 1.73018312 55.06595682
2 RAD 874.1452451507120 TH 1.00000000 AIR
3 RAD 43.3723238775944 TH 7.22734397
3 GLM 1.90000000 31.85177396
4 RAD 75.5982585496480 TH 3.72513003 AIR
5 RAD 136.0674307510774 TH 4.22284549
5 GLM 1.67482513 42.21779395
6 RAD 21.1183776902342 TH 9.49462964 AIR
7 RAD 50.9917098403988 TH 15.50072204
7 GLM 1.85120210 23.62968294
8 RAD 27.2208332948113 TH 9.61128996 AIR
9 RAD 105.6629879994650 TH 21.75575719
9 GLM 1.50100000 77.45538462
10 RAD -27.0616542071280 TH 38.57499099 AIR
11 RAD -348.1756599714766 TH 5.06911986
11 GLM 1.68456543 59.52245393
12 RAD -51.1885698039190 TH 2.16493310 AIR
13 RAD -40.9131894386343 TH 25.47601346
13 GLM 1.55595432 44.66796234
14 RAD -87.3681162352767 TH 50.80907300 AIR
14 CV -0.01144582
14 UMC -0.13986014
14 TH 50.80907300
14 YMT 0.00000000
15 CAO 32.00000000 0.00000000 0.00000000
15 CV 0.00000000000000 TH 0.00000000 AIR
ZFILE 1
CAM RANK 2
CAM EXPONENT 1.00000
14 14
ZOOM 2

```

```

OBA 1000.00000 -366.554000 12.7000 0.0000 0.00000000 0.0000 12.7000
ZDATA 0.0000000E+00
END

```



So that's how it's done: Figure out what's wrong and use the tools in SYNOPSIS to fix it. Sometimes it's quick and sometimes not. That's what lens design is all about, blind alleys and all.

But that is probably enough for this lesson.

Oh, we almost forgot: Why did we enter the surface number (14) for the zooming group, since the YMT solve will override it anyway? Well, the program requires a group definition, and it won't work otherwise. That's to save you from a serious mistake if you ever leave those data out for a real zoom lens.

We will revisit this problem in Lesson 17 and show how yet other tools can be effectively applied and will save some time.